

# Wonders of $E_{10}$ and $K(E_{10})$

Hermann Nicolai

MPI für Gravitationsphysik (AEI), Potsdam

*Wonders of Gauge Theory and Supergravity*

IHP and LPT-ENS, Paris, 23 - 28 June 2008

(mostly) based on work done in collaboration with:

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Main message of this talk:

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... and perhaps also for quantum gravity...

# The BKL Paradigm

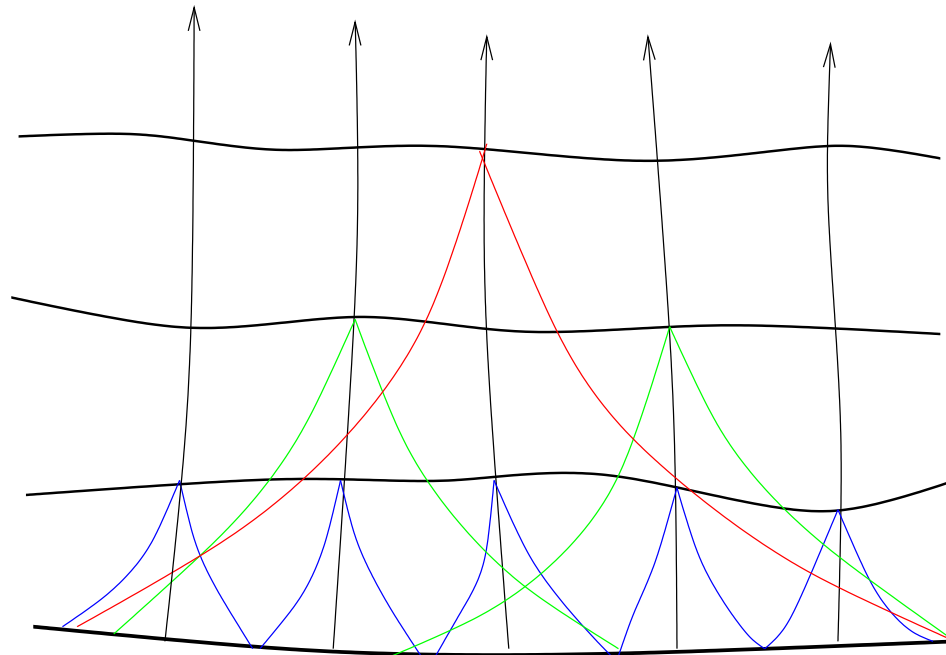
- Near a spacelike (cosmological) singularity, Einstein equations should simplify  $\Rightarrow$  BKL decoupling:  $\partial_x \ll \partial_t$ ?

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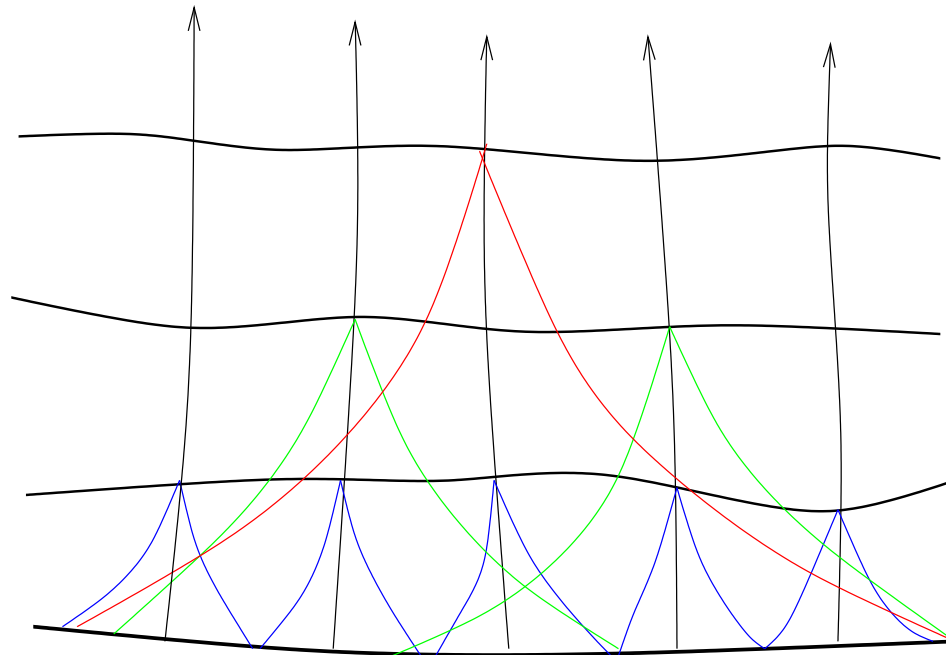
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- Dimensional reduction to one (time) dimension  $\rightarrow$  **effective dynamics** near singularity from **gradient expansion**?  $\rightarrow$  billiards, chaotic oscillations, etc.

# Another (old) paradigm

- Cosmological evolution as ‘geodesic motion’ in the moduli space of 3-geometries [Wheeler, DeWitt,...]:

$$\mathcal{M} \equiv \mathcal{G}^{(3)} = \frac{\{\text{spatial metrics } g_{ij}(\mathbf{x})\}}{\{\text{diffeomorphisms}\}}$$

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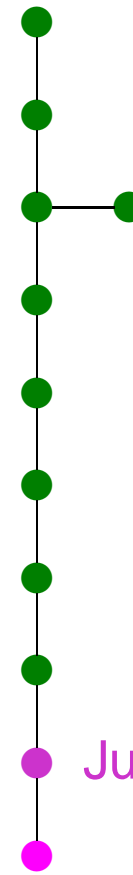
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- Unification of space-time, matter and gravitation: configuration space  $\mathcal{M}$  for quantum gravity should consistently incorporate matter degrees of freedom.

# Hidden symmetries

Reduction of  $\text{SUGRA}_{11}$  to  $D = 11 - n$  [Cremmer, Julia (1979)]

$n$	Scalar Coset $E_n/K(E_n)$
1	$GL(1)/\mathbf{1}$
2	$GL(2)/SO(2)$
3	$SL(3) \times SL(2)/U(2)$
4	$SL(5)/SO(5)$
5	$SO(5, 5)/SO(5) \times SO(5)$
6	$E_6/USp(4)$
7	$E_7/SU(8)$
8	$E_8/(Spin(16)/\mathbb{Z}_2)$
9	$E_9/K(E_9)$
10	$E_{10}/K(E_{10})$
11	$E_{11}/K(E_{11})$



Julia('85), DHN('02)

West (2001)



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However:  $\mathcal{L} = \mathcal{L}(g_{ij}(t), A_{ijk}(t))$  is only invariant under  $GL(10, \mathbb{R}) \times T_{120} \dots$  **but:**

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Effective dynamics of diagonal metric degrees of freedom is governed by *cosmological billiards* in Weyl chamber of  $E_{10}$ !

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motivates **BASIC CONJECTURE:**  $\mathcal{M} = E_{10}/K(E_{10})$

Dynamics of supergravity (or some M theoretic extension)

Null geodesic motion on  $E_{10}/K(E_{10})$  coset space

are **equivalent!** [DHN, hep-th/0207267]

**SUGRA eqs. of motion + canonical constraints**

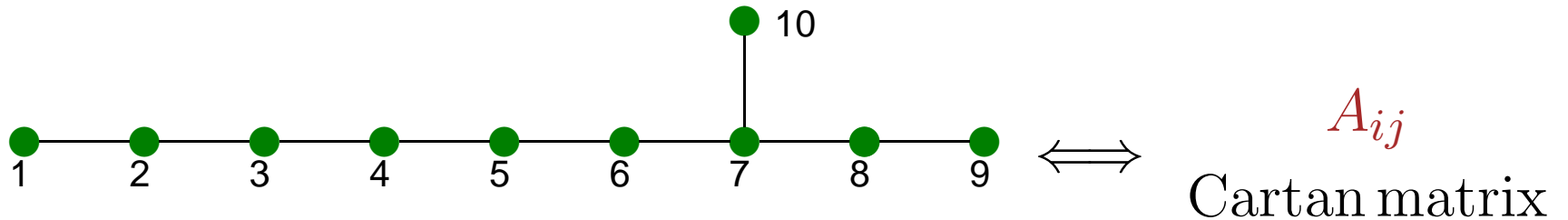
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# Definition of $E_{10}$

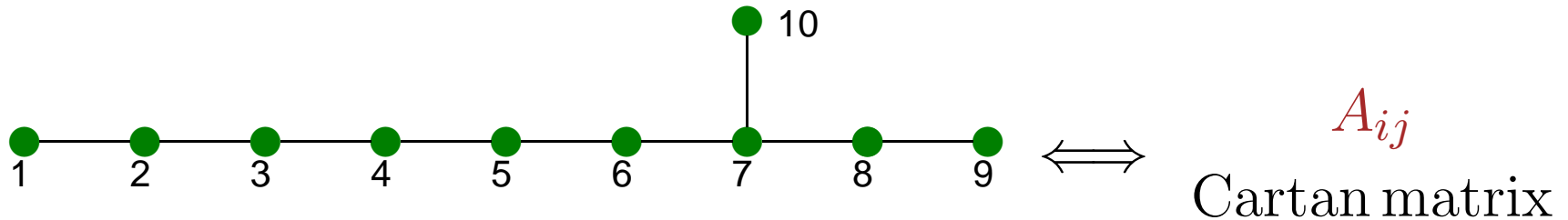
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**Chevalley–Serre presentation:** Generators  $h_i, e_i, f_i$  for  $i = 1, \dots, 10$  with relations

$$[h_i, h_j] = 0,$$

$$[e_i, f_j] = \delta_{ij} h_i,$$

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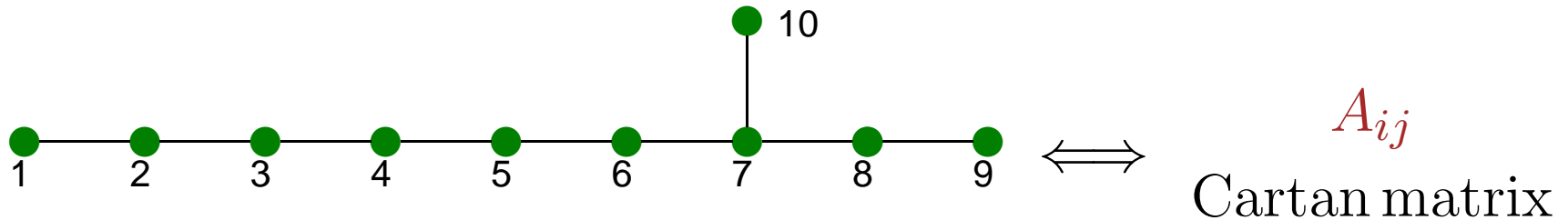
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$h_i$  span Cartan subalgebra  $\mathfrak{h}$ ;  $e_i$  and  $f_i$ : positive and negative simple root generators

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- Triangular decomposition  $\rightarrow$  **Computability**

$$\mathfrak{g} = \mathfrak{e}_{10} = \mathfrak{n}_- \oplus \mathfrak{h} \oplus \mathfrak{n}_+ \quad , \quad \text{with } \mathfrak{n}_\pm := \bigoplus_{\alpha \gtrless 0} \mathfrak{g}_\alpha$$

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$$\mathfrak{k}_{10} \equiv K(\mathfrak{e}_{10}) = \{x \in \mathfrak{e}_{10} : \omega(x) \equiv -x^T = x\}$$

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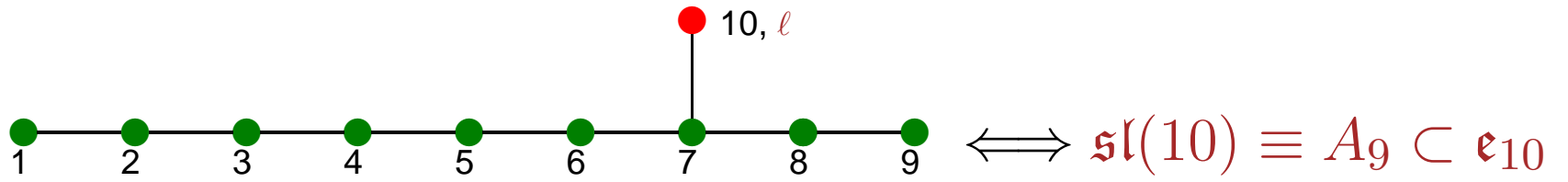
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**However:**  $\mathfrak{k}_{10}$  is *not* a Kac-Moody algebra [KN, hep-th/0506238]

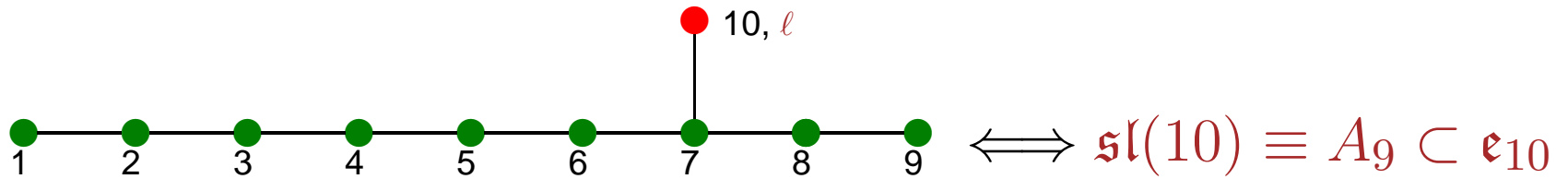
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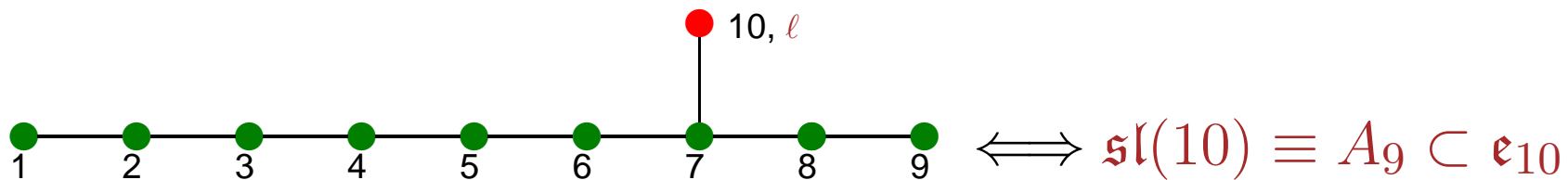


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$\ell$	$A_9$ module	Tensor
0	$[1000000001] \oplus [0000000000]$	$K^a_b$
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These are just the representations corresponding to the bosonic fields of  $D = 11$  SUGRA and their magnetic duals.

At level  $\ell = 3$ : dual graviton  $h_{a_1 \dots a_8 | a_9}$  (with  $h_{[a_1 \dots a_8 | a_9]} = 0$ )

[For more representations, see: Fischbacher, N. hep-th/0301017]

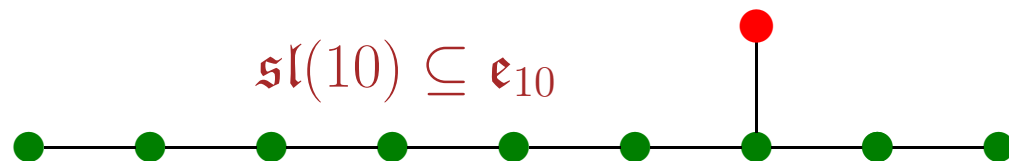
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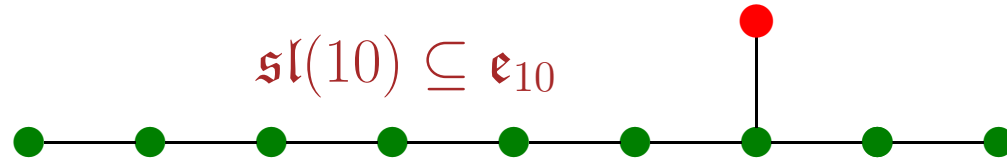


$D = 11$  SUGRA

[DHN; West 2002]

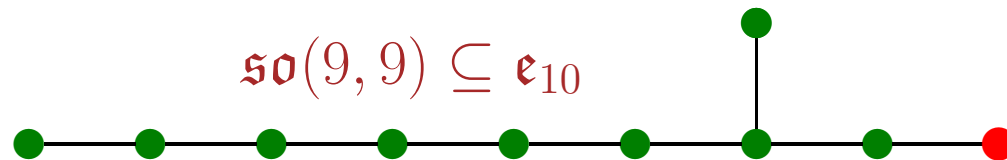
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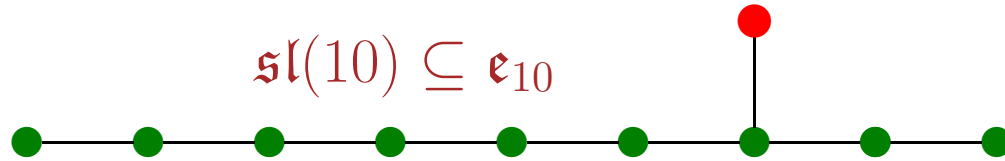
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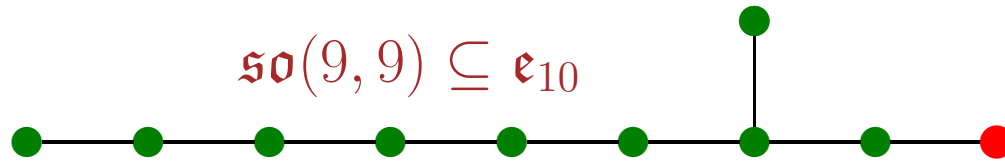
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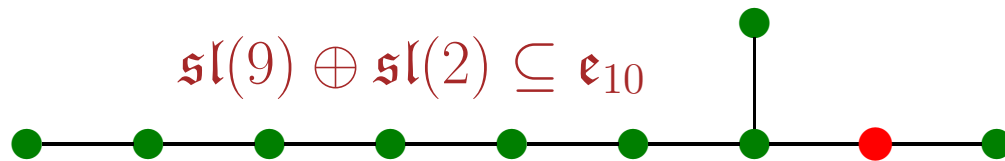
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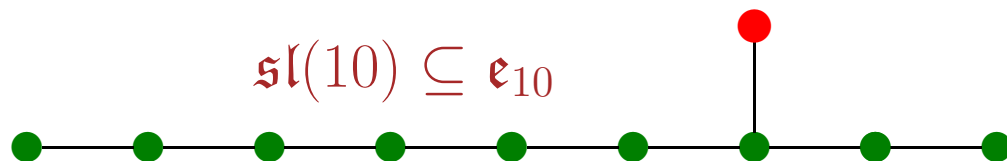
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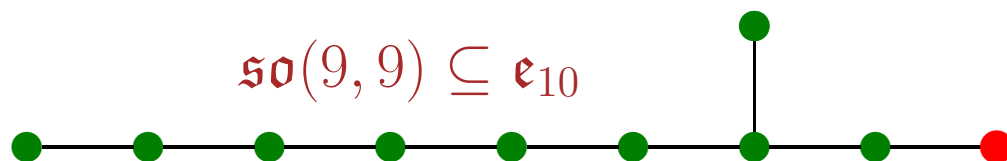
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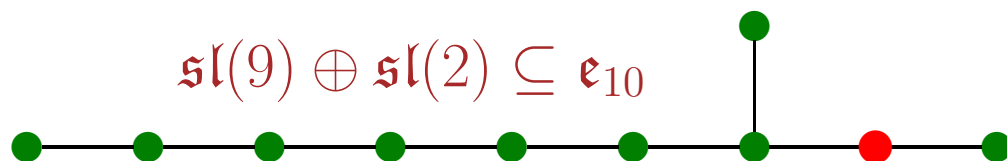


$$\mathfrak{so}(9, 9) \subseteq \mathfrak{e}_{10}$$

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$$\mathfrak{sl}(9) \oplus \mathfrak{sl}(2) \subseteq \mathfrak{e}_{10}$$

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These are the (maximal) low energy theories of the ‘M-theory diagram’, now all part of a single model.



# Dynamics: bosonic Lagrangian

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Equations of motion: **null geodesic** on  $E_{10}/K(E_{10})$

$$n \partial_t (n^{-1} \mathcal{P}) = [\mathcal{Q}, \mathcal{P}] \quad , \quad \langle \mathcal{P} | \mathcal{P} \rangle = 0.$$

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With  $\partial_t \mathcal{V} \mathcal{V}^{-1} = \sum_{\ell \geq 0} P^{(\ell)} * E^{(\ell)}$  (schematically) and truncation  $P^{(\ell)} = 0$  for  $\ell > 3 \Rightarrow$

Equations of motion up to  $\ell = 3$  ( $a, b = 1, \dots, 10$ ) [DHN; DN, hep-th/0410245]

$$\begin{aligned}
 n\mathcal{D}^{(0)}(n^{-1}P_{ab}^{(0)}) &= -\frac{1}{4}(P_{acd}^{(1)}P_{bcd}^{(1)} - \frac{1}{9}\delta_{ab}P_{cde}^{(1)}P_{cde}^{(1)}) \\
 &\quad -\frac{1}{2 \cdot 5!}(P_{ac_1\dots c_5}^{(2)}P_{bc_1\dots c_5}^{(2)} - \frac{1}{9}\delta_{ab}P_{c_1\dots c_6}^{(2)}P_{c_1\dots c_6}^{(2)}) \\
 &\quad +\frac{4}{9!}(P_{ac_1\dots c_7|c_8}^{(3)}P_{bc_1\dots c_7|c_8}^{(3)} + \frac{1}{8}P_{c_1\dots c_8|a}^{(3)}P_{c_1\dots c_8|b}^{(3)} \\
 &\quad \quad -\frac{1}{8}\delta_{ab}P_{c_1\dots c_8|c_9}^{(3)}P_{c_1\dots c_8|c_9}^{(3)}) \\
 n\mathcal{D}^{(0)}(n^{-1}P_{abc}^{(1)}) &= -\frac{1}{6}P_{abcdef}^{(2)}P_{def}^{(1)} + \frac{1}{3 \cdot 5!}P_{abcd_1\dots d_5|d_6}^{(3)}P_{d_1\dots d_6}^{(2)} \\
 n\mathcal{D}^{(0)}(n^{-1}P_{a_1\dots a_6}^{(2)}) &= \frac{1}{6}P_{a_1\dots a_6cde}^{(3)}P_{cde}^{(1)} \\
 n\mathcal{D}^{(0)}(n^{-1}P_{a_1\dots a_8|a_9}^{(3)}) &= 0 \quad (\text{with } P_{[a_1\dots a_8|a_9]}^{(3)} = 0).
 \end{aligned}$$

This is a **consistent truncation** of  $E_{10}/K(E_{10})$  coset dynamics: solutions of truncated theory are also solutions of the full theory.

# Correspondence with SUGRA<sub>11</sub>



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Bosonic  $D = 11$  supergravity equations [Cremmer, Julia, Scherk 1978]

$$\mathcal{E}_{AB} \equiv R_{AB} - \frac{1}{3}F_{ACDE}F_B{}^{CDE} + \frac{1}{36}\eta_{AB}F_{CDEF}F^{CDEF} = 0$$

$$\mathcal{M}^{BCD} \equiv D_A F^{ABCD} + \frac{1}{576}\epsilon^{BCDE_1\dots E_8}F_{E_1\dots E_4}F_{E_5\dots E_8} = 0$$

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Consider gauge fixed (*à la* ADM) equations *at some fixed spatial point*  $\mathbf{x}_0$ :

- keeping all temporal and first order spatial derivatives at  $\mathbf{x}_0$
- zero-shift gauge:  $E_M{}^A = \left( \begin{array}{c|c} N & 0 \\ \hline 0 & e_m{}^a \end{array} \right)$  and Coulomb gauge:  $A_{tmn} = 0$
- Anholonomy coefficients  $[\partial_b, \partial_c] = \tilde{\Omega}_{bc|a}\partial_a$  chosen traceless (in some neighborhood of  $\mathbf{x}_0$ ) by exploiting *spatial* Lorentz group, i.e.  $\Lambda_{ab} = \Lambda_{ab}(t, \mathbf{x})$  [???
- Thus the standard ADM procedure leads to usual split into:
  - Dynamical equations:  $\mathcal{E}_{ab} = \mathcal{M}_{abc} = D_{[0}F_{bcde]} = R_{[0a b]c} = 0$
  - Canonical constraints:  $\mathcal{E}_{00} = \mathcal{E}_{0a} = \mathcal{M}_{0ab} = D_{[a}F_{bcde]} = R_{[ab c]d} = 0$

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Then with the identification  $n = Ne^{-1}$  and (r.h.s. always at fixed spatial point  $\mathbf{x} = \mathbf{x}_0$ )

$$\begin{aligned}\mathcal{D}^{(0)}P_{ab}^{(0)} &= R_{ab}^{\text{time derivatives}} \\ P_{abc}^{(1)} &= NF_{0abc} \\ P_{a_1\dots a_6}^{(2)} &= -\frac{1}{4!}N\epsilon_{a_1\dots a_6 b_1\dots b_4}F_{b_1\dots b_4} \\ P_{a_1\dots a_8|a_9}^{(3)} &= \frac{3}{2}N\epsilon_{a_1\dots a_8 bc}\tilde{\Omega}_{bc|a_9}\end{aligned}$$

the two sets of dynamical equations coincide! (recall  $P_{[a_1\dots a_8|a_9]}^{(3)} = 0 \Leftrightarrow \tilde{\Omega}_{ab|b} = 0$ )

Dynamical equations for **mIIA** and **IIB** similarly from level decompositions w.r.t. finite dimensional subgroups  $D_9 \equiv SO(9, 9) \subset E_{10}$  and  $A_8 \times A_1 \equiv SL(9) \times SL(2) \subset E_{10}$ .

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Conserved  $E_{10}$  current  $\mathcal{J} = n\mathcal{V}\mathcal{P}\mathcal{V}^{-1}$  ( $\equiv$  Noether charge associated with global  $E_{10}$ ):

$$\begin{aligned}\mathcal{J} = & \frac{1}{9!} J_{(-3)}^{m_0|m_1\dots m_8} F_{m_0|m_1\dots m_8} + \frac{1}{6!} J_{(-2)}^{m_1\dots m_6} F_{m_1\dots m_6} + \frac{1}{3!} J_{(-1)}^{mnp} F_{mnp} \\ & + J_{(0)m}^n K^m_n + \frac{1}{3!} J_{(1)mnp} E^{mnp} + \frac{1}{6!} J_{(2)m_1\dots m_6} E^{m_1\dots m_6} + \dots\end{aligned}$$

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Consider Sugawara-like ( $\propto \mathcal{J} \otimes \mathcal{J}$ ) expressions [DKN, hep-th 0709.2691]

$$\begin{aligned} \mathfrak{L}_{(-6)}^{m_1\dots m_{10}; n_0|n_1\dots n_7} &= J_{(-3)}^{n_0|m_1\dots m_8} J_{(-3)}^{m_9|m_{10}n_1\dots n_7} \\ \mathfrak{L}_{(-5)}^{m_1\dots m_{10}; n_1\dots n_5} &= J_{(-2)}^{n_1\dots n_4 m_1 m_2} J_{(-3)}^{m_3\dots m_{10}} \\ \mathfrak{L}_{(-4)}^{m_1\dots m_{10}; n_1 n_2} &= \frac{21}{5} J_{(-2)}^{n_1 m_1\dots m_5} J_{(-2)}^{n_2 m_6\dots m_{10}} + J_{(-1)}^{n_1 m_1 m_2} J_{(-3)}^{n_2|m_3\dots m_{10}} \end{aligned}$$

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(with appropriate antisymmetrizations) to re-express canonical constraints:

$$\begin{aligned} \mathfrak{L}_{(-6)}^{m_1\dots m_{10}; n_0|n_1\dots n_7} &\propto \epsilon^{m_1\dots m_{10}} \epsilon^{n_1\dots n_7 pqr} R_{pq}{}_{rn_0} && \text{Bianchi (I)} \\ \mathfrak{L}_{(-5)}^{m_1\dots m_{10}; n_1\dots n_5} &\propto \epsilon^{m_1\dots m_{10}} \epsilon^{n_1\dots n_5 p_1\dots p_5} D_{p_1} F_{p_2\dots p_5} && \text{Bianchi (II)} \\ \mathfrak{L}_{(-4)}^{m_1\dots m_{10}; n_1 n_2} &\propto \epsilon^{m_1\dots m_{10}} \mathcal{M}^{0n_1 n_2} && \text{Gauss constraint} \\ \mathfrak{L}_{(-3)}^{m_1\dots m_9} &\propto \epsilon^{m_1\dots m_9 n} \mathcal{E}_{0n} && \text{Momentum constraint} \end{aligned}$$

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Thank you for your attention